# Structure Factor Algebra in the Probabilistic Procedure for Phase Determination. I 

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It is possible to take the statistical weight of reflexions into account in the $\Sigma_{1}$, Sayre and tangent formulae. A suitable new use of the normalized structure factors is proposed in these procedures for phase determination; new generalized formulae are derived in a form suitable for automatic computer calculation.

## Introduction

Hauptman \& Karle (1953) defined the normalized structure factor $E_{\mathrm{h}}$ as

$$
\begin{equation*}
E_{\mathbf{h}}^{2}=F_{\mathbf{h}}^{2} / \varepsilon \sum_{1}^{N} f_{j}^{2}(\mathbf{h}) \tag{1}
\end{equation*}
$$

where
$\varepsilon=\frac{m_{20}+m_{02}}{m} ; m_{20}=\iiint \psi^{2}(\mathbf{h}) \mathrm{d} \mathbf{r} ; m_{02}=\iiint \eta^{2}(\mathbf{h}) \mathrm{d} \mathbf{r}$.
$m$ is the symmetry number of the space group, $\psi$ and $\eta$ are the trigonometric functions for the real and imaginary parts of the structure factor.
The quasi-normalized structure factor $\mathscr{E}_{\mathrm{h}}$ is also frequently used in structure determination: it is

$$
\begin{equation*}
\mathscr{E}_{\mathrm{h}}=\sigma_{2}^{-1 / 2} \sum_{1}^{N} Z_{j} \exp 2 \pi i \mathrm{kr}_{j} \tag{2}
\end{equation*}
$$

where $\sigma_{2}=\sum_{1}^{N} Z_{j}^{2}$ : no space-group weight of reflexions is considered. The $E_{\mathbf{h}}$ was defined to ensure always that mean-square $\left\langle E_{\mathrm{h}}^{2}\right\rangle=1$, the quasi-normalized $\mathscr{E}_{\mathrm{h}}$ to guarantee simplicity in the derivation of algebraic relationships.
Karle \& Karle (1966) advised, on an experimental basis, the use in symbolic-addition procedures of the $E_{\mathrm{h}}^{\prime}$ factors where

$$
\begin{equation*}
\left|E_{\mathbf{h}}^{\prime}\right|^{2}=\frac{\left|F_{\mathbf{h}}\right|^{2}}{\varepsilon^{\prime} \sum_{j}^{N} f_{j}^{2}(\mathbf{h})} . \tag{3}
\end{equation*}
$$

$\varepsilon^{\prime}$ is a number which corrects for space-group extinctions: the relation proposed between $\mathscr{E}_{\mathrm{h}}$ and $E_{\mathrm{h}}^{\prime}$ is

$$
\left|\mathscr{E}_{\mathbf{h}}^{\circ}\right|^{2}(1-q)=\left|E_{\mathbf{h}}^{\prime}\right|^{2},
$$

where $q$ is the fraction of reflexions in the $\mathbf{h}$ set which are space-group extinctions.
This work justifies in Sayre, tangent and $\Sigma_{1}$ formulas a new use of the normalized structure factors on the basis of their algebra; a combination of the linearization theory (Bertaut \& Waser, 1957; Bertaut, 1959a,b) and of the probability distribution functions have been used to derive the method in a form suitable for automatic computing. A method, similar in some
aspects, has been used in the centrosymmetric case by Naya, Nitta \& Oda (1964).

## Algebraic considerations

For equal atoms, if $m$ is the space-group order, according to (2), (Bertaut \& Waser, 1957)

$$
\begin{align*}
\mathscr{E}_{\mathbf{h}}=N^{-1 / 2} \sum_{1}^{N / m} p_{\mathbf{h}} \sum_{1}^{m / p \mathbf{h}} \exp 2 \pi i \mathbf{h C} \mathbf{C}_{s} \mathbf{x}_{j} & \\
& =N^{-1 / 2} \sum_{1}^{N / m} \xi_{j}(\mathbf{h}) \tag{4}
\end{align*}
$$

where $\mathbf{C}_{s} \mathbf{x}=\mathbf{R}_{s} \mathbf{x}+\mathbf{T}_{s}: \mathbf{C}_{s}$ is the $s$-symmetry operation ( $\mathbf{R}_{s}$ rotation component, $\mathbf{T}_{s}$ translation component), $p_{\mathrm{h}}$ is the statistical weight.

Following Woolfson (1954), we impose for general reflexions the condition that $\mathbf{k}$ varies while $\mathscr{E}_{\mathbf{k}}$ and $\mathscr{E}_{\mathrm{h}-\mathrm{k}}$ are constant: we obtain, by the application of the central-limit theorem (Cramér, 1951).

$$
\begin{align*}
& \left\langle\mathscr{E}_{\mathbf{h}}\right\rangle=N^{-1 / 2} \sum_{1}^{N / m}\left\langle\xi_{j}(\mathbf{h})\right\rangle=N^{-1 / 2} \mathscr{E}_{\mathbf{k}} \mathscr{E}_{\mathbf{h}-\mathbf{k}}  \tag{5}\\
& \begin{aligned}
V_{\mathbf{h}}=N^{-1} \sum_{1}^{N / m} & \left.\left.\langle | \xi_{j}(h)\right|^{2}\right\rangle-\left|\left\langle\xi_{j}(h)\right\rangle\right|^{2} \\
& =N^{-1} \sum_{1}^{N / m}\left\{m-\frac{\left|\mathscr{E}_{\mathbf{k}} \mathscr{E}_{\mathbf{h}-\mathbf{k}}\right|^{2} m^{2}}{N^{2}}\right\} .
\end{aligned}
\end{align*}
$$

If the number $N / m$ of the independent atoms in the cell is large enough

$$
\begin{equation*}
V_{\mathrm{h}}=1 . \tag{7}
\end{equation*}
$$

By following the Cochran (1955) treatment and expressing in $E$ terms, we obtain the well known results:

$$
\begin{gather*}
\left\langle\varphi_{\mathbf{h}}\right\rangle=\varphi_{\mathbf{k}}+\varphi_{\mathbf{h}-\mathbf{k}},  \tag{8}\\
P\left(\varphi_{\mathbf{h}}\right)=\exp \left\{G_{\mathbf{h}, \mathbf{k}} \cos \left(\varphi_{\mathbf{h}}-\varphi_{\mathbf{k}}-\varphi_{\mathbf{h}-\mathbf{k}}\right)\right\} / \\
2 \pi I_{0}\left(G_{\mathbf{h}, \mathbf{k}}\right), \tag{9}
\end{gather*}
$$

where
$G_{\mathbf{h}, \mathbf{k}}=2 \frac{\left|E_{\mathbf{h}} E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\right|}{V N}$,
and $I_{0}$ is a modified Bessel function of the second kind (Watson, 1922).

Starting from equations (5), (7), (8) and (9) Karle \& Karle (1966) established, by probability considerations, the tangent formula

$$
\begin{equation*}
\tan \varphi_{\mathbf{h}}=\tan \frac{\sum_{\mathbf{k}}\left|E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\right| \sin \left(\varphi_{\mathbf{k}}+\varphi_{\mathbf{h}-\mathbf{k}}\right)}{\sum_{\mathbf{k}}\left|E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\right| \cos \left(\varphi_{\mathbf{k}}+\varphi_{\mathbf{h}-\mathbf{k}}\right)} \tag{10}
\end{equation*}
$$

with variance

$$
\begin{equation*}
V_{\mathrm{h}}=\frac{\pi^{2}}{3}+\left[I_{0}(\alpha)\right]^{-1} \sum_{1}^{\infty} \frac{I_{2 n}(\alpha)}{n^{2}}-4\left[I_{0}(\alpha)\right]^{-1} \sum_{0}^{\infty} \frac{I_{2 n+1}(\alpha)}{(2 n+1)^{2}} \tag{11}
\end{equation*}
$$

where

$$
\begin{align*}
\alpha= & \left\{\left[\sum_{\mathbf{k}} G_{\mathbf{h}, \mathbf{k}} \cos \left(\varphi_{\mathbf{k}}+\varphi_{\mathbf{h}-\mathbf{k}}\right)\right]^{2}\right. \\
& \left.+\left[\sum_{\mathbf{k}} G_{\mathbf{h}, \mathbf{k}} \sin \left(\varphi_{\mathbf{k}}+\varphi_{\mathbf{h}-\mathbf{k}}\right)\right]^{2}\right\}^{1 / 2} \tag{12}
\end{align*}
$$

The central-limit theorem, employed to obtain equations (5), (7), (8), (9) disregards the actual algebraic form of the $\xi$ function, and therefore these equations are not strictly applicable in the case of special reflexions.

In order to generalize the previous formulae probability theory will be used here.

## Probability considerations

Following Hauptman \& Karle (1953) and employing the Klug (1958) notation, for a general centrosymmetrical group of order $m$, the joint probability distribution results:

$$
\begin{align*}
& P\left(E_{1}, E_{2}, E_{3}\right)=\frac{1}{(2 \pi)^{3 / 2}} \exp \left[-\frac{1}{2}\left(E_{1}^{2}+E_{2}^{2}+E_{3}^{2}\right)\right] \\
& \quad \times\left\{1+\frac{1}{t^{1 / 2}}\left[\frac{\lambda_{111}}{1!1!1!} E_{1} E_{2} E_{3}\right]\right. \\
& \quad+\frac{1}{t}\left[\frac{\lambda_{400}}{4!0!0!} H_{4}\left(E_{1}\right)+\frac{\lambda_{040}}{0!4!0!} H_{4}\left(E_{2}\right)+\ldots\right] \\
& \quad+\frac{1}{2 t}\left[\frac{\lambda^{2} 111}{1!1!1!} H_{2}\left(E_{1}\right) H_{2}\left(E_{2}\right) H_{2}\left(E_{3}\right)\right] \\
& \quad+\frac{1}{t^{3 / 2}}\left[\frac{\lambda_{113}}{1!1!3!} H_{1}\left(E_{1}\right) H_{1}\left(E_{2}\right) H_{3}\left(E_{3}\right)+\ldots\right] \tag{13}
\end{align*}
$$

where

$$
E_{1}=E_{\mathbf{h}}, E_{2}=E_{\mathbf{k}}, E_{3}=E_{\mathbf{h}+\mathbf{k}},
$$

and

$$
\lambda_{r s w}=\frac{K_{r s w}}{K_{200}^{r / 2} K_{020}^{(T, 2} K_{002}^{w / 2}}=\frac{K_{r s w}}{(m)^{(r+s+w) / 2}} .
$$

$K_{r s w}$ is a multivariate cumulant of order $r+s+w$, and $H(z)$ is a Hermite polynomial defined by the equation:

$$
H_{v}(x)=(-1)^{v} \exp \left[\frac{1}{2} x^{2}\right] \frac{\mathrm{d}^{v}}{\mathrm{~d} x^{v}} \exp \left[-\frac{1}{2} x^{2}\right]
$$

The first moment of the conditional probability distribution $P\left(E_{1} \mid E_{2}, E_{3}\right)$ is, retaining terms to order $1 / t^{1 / 2}$,

$$
\begin{equation*}
\left\langle E_{1} \mid E_{2}, E_{3}\right\rangle=\frac{\lambda_{111}}{t^{1 / 2}} E_{2} E_{3} \tag{14}
\end{equation*}
$$

As

$$
K_{111}=m_{111}=\left\langle\frac{\xi(\mathbf{h})}{\sqrt{p_{\mathbf{h}}}} \frac{\xi(\mathbf{k})}{\sqrt{p_{\mathbf{k}}}} \frac{\xi(\mathbf{h}+\mathbf{k})}{\sqrt{p_{\mathbf{h}+\mathbf{k}}}}\right\rangle
$$

where $p_{\mathbf{h}}, p_{\mathbf{k}}, p_{\mathrm{h}+\mathbf{k}}$ are the statistical weights of $E_{\mathbf{h}}$, $E_{\mathbf{k}}, E_{\mathbf{h}+\mathbf{k}}$, we find

$$
\begin{align*}
& \left\langle E_{\mathbf{h}} \mid E_{\mathbf{k}}, E_{\mathbf{h}+\mathbf{k}}\right\rangle=\left\langle\frac{\xi(\mathbf{h}) \xi(\mathbf{k}) \xi(\mathbf{h}+\mathbf{k})}{m^{3 / 2} \sqrt{p_{\mathbf{h}} p_{\mathbf{k}} p_{\mathbf{h}+\mathbf{k}}}}\right\rangle \\
& \quad \times \frac{1}{t^{1 / 2}} E_{\mathbf{k}} E_{\mathbf{h}+\mathbf{k}}=W_{\mathbf{h}, \mathbf{k}}\left(N^{-1 / 2} E_{\mathbf{k}} E_{\mathbf{h}+\mathbf{k}}\right), \tag{15}
\end{align*}
$$

where

$$
\begin{equation*}
W_{\mathbf{h}, \mathbf{k}}=\frac{1}{m \sqrt{p_{\mathbf{h}} p_{\mathbf{k}} p_{\mathbf{h}+\mathbf{k}}}}\left\langle\sum_{1}^{m}, r \xi\left[\mathbf{h}\left(\mathbf{C}_{s}-\mathbf{I}\right)+\mathbf{k}\left(\mathbf{C}_{r}-\mathbf{I}\right)\right]\right\rangle . \tag{16}
\end{equation*}
$$

$W_{\mathrm{h}, \mathbf{k}}$ takes the statistical weights of the normalized structure factors $E_{\mathbf{h}}, E_{\mathbf{k}}, E_{\mathbf{h}+\mathbf{k}}$ into account. Formula (16) has been worked out in the Appendix and is very suitable for automatic computing.

As is well known, the second moment of the $E_{\mathrm{h}}$ conditional distribution is, from equation (13), retaining terms to order $1 / t^{1 / 2}$, equal to unity, whatever the statistical weights may be.

If we expand the $E_{\mathbf{h}}$ conditional probability in the form of the Gram-Charlier series (Cramér, 1951) we obtain

$$
\begin{aligned}
& P\left(E_{\mathbf{h}} \mid E_{\mathbf{k}}, E_{\mathbf{h}+\mathbf{k}}\right)=\frac{1}{\sqrt{2 \pi}} \\
& \quad \times \exp \left[-\frac{1}{2}\left(E_{\mathbf{h}}-\frac{W_{\mathbf{h}, \mathbf{k}}}{N^{1 / 2}} E_{\mathbf{k}} E_{\mathbf{h}+\mathbf{k}}\right)^{2}\right]+\ldots
\end{aligned}
$$

As

$$
P_{+}=\left(\frac{P_{-}}{P_{+}}+1\right)^{-1}
$$

we easily obtain

$$
P_{+}\left(E_{\mathbf{h}}\right)=\frac{1}{2}+\frac{1}{2} \tanh \left[\frac{W_{\mathbf{h}, \mathbf{k}}}{N^{1 / 2}}\left|E_{\mathbf{h}}\right| E_{\mathbf{k}} E_{\mathbf{h}+\mathbf{k}}\right]
$$

or in general

$$
\begin{equation*}
P_{+}\left(E_{\mathbf{h}}\right)=\frac{1}{2}+\frac{1}{2} \tanh \left[\frac{\left|E_{\mathrm{h}}\right|}{N^{1 / 2}} \sum_{1}^{r} W_{\mathbf{h}, \mathbf{k} j} E_{\mathbf{k} j} E_{\mathbf{h}+\mathbf{k}_{j}}\right] . \tag{17}
\end{equation*}
$$

## $\boldsymbol{\Sigma}_{1}$ formula

From Klug (1958) we derive, for a centrosymmetric space group of order $m$, the probability distribution

$$
\begin{align*}
P\left(E_{1}, E_{2}\right) & =\frac{1}{2 \pi} \exp \left[-\frac{1}{2}\left(E_{1}^{2}+E_{2}^{2}\right)\right] \\
& \times\left\{1+\frac{1}{t^{1 / 2}} \frac{\lambda_{12}}{1!2!} H_{1}\left(E_{1}\right) H_{2}\left(E_{2}\right)\right. \\
& +\frac{1}{t}\left[\frac{\lambda_{40}}{4!0!} H_{4}\left(E_{1}\right)+\frac{\lambda_{04}}{0!4!} H_{4}\left(E_{2}\right)\right. \\
& \left.\left.+\frac{1}{2}\left(\frac{\lambda_{12}}{1!2!}\right)^{2} H_{2}\left(E_{1}\right) H_{4}\left(E_{2}\right)\right]+\ldots\right\} \tag{18}
\end{align*}
$$

where

$$
E_{1}=E_{2 \mathrm{~h}}, \quad E_{2}=E_{\mathrm{h}}
$$

and

$$
\lambda_{i j}=\frac{K_{i j}}{\left(K_{20}\right)^{i / 2}\left(K_{02}\right)^{j / 2}}=\frac{K_{i j}}{m^{(i+j) / 2}}
$$

The first conditional moment $\left\langle E_{2 h} \mid E_{\mathrm{h}}\right\rangle$ gives, retaining terms to order $1 / t^{1 / 2}$,

$$
\left\langle E_{2 \mathbf{h}} \mid E_{\mathbf{h}}\right\rangle=\frac{1}{t^{1 / 2}} \frac{\lambda_{12}}{1!2!}\left(E_{\mathbf{h}}^{2}-1\right)
$$

As (see Appendix)

$$
\begin{aligned}
K_{12}=m_{12} & =\frac{\left\langle\xi^{2}(\mathbf{h}) \xi(2 \mathbf{h})\right\rangle}{\left(p_{\mathbf{h}}\right)^{3 / 2}} \\
& =\left\langle\frac{\sum_{q}^{m} \xi\left[\mathbf{h}\left(\mathbf{I}-\mathbf{R}_{q}\right)\right] \xi\left[\mathbf{h}\left(\mathbf{I}-\mathbf{R}_{2}\right)\right]}{p_{\mathbf{h}}^{3 / 2}}\right\rangle=\frac{m}{\sqrt{p_{\mathbf{h}}}}
\end{aligned}
$$

we obtain

$$
\left\langle E_{2 \mathrm{~h}} \mid E_{\mathrm{h}}\right\rangle=\frac{1}{2 N^{1 / 2} \sqrt{p_{\mathrm{h}}}}\left(E_{\mathrm{h}}^{2}-1\right)
$$

As the variance is equal to unity, by expanding the $E_{2 h}$ conditional distribution in Gram-Charlier series, we can write

$$
P\left(E_{2 \mathrm{~h}} \mid E_{\mathrm{h}}\right)=\frac{1}{\sqrt{2 \pi}} \exp \left[-\frac{1}{2}\left(E_{2 \mathrm{~h}}-\frac{\left(E_{\mathrm{h}}^{2}-1\right)}{2 N^{1 / 2} \sqrt{p_{\mathrm{h}}}}\right)^{2}\right]
$$

This equation can be compared with previous results [i.e. Cochran \& Woolfson (1955), equation (3.8)].

The probability $P_{+}\left(E_{2 \mathbf{h}}\right)$ is finally obtained as

$$
\begin{equation*}
P_{+}\left(E_{2 \mathrm{~h}}\right)=\frac{1}{2}+\frac{1}{2} \tanh \left[\frac{1}{2 N^{1 / 2}} \frac{\sqrt{p_{\mathbf{h}}}}{}\left|E_{2 \mathbf{h}}\right|\left(E_{\mathbf{h}}^{2}-1\right)\right] \tag{19}
\end{equation*}
$$

To obtain other $\sum_{1}$ formulas, we modify the probability distribution (18) by putting $E_{2}=E_{\mathbf{h}}\left(\mathbf{I}-\mathbf{R}_{s}\right)$, where $\mathbf{R}_{s}$ is a matrix rotation of the space groups. In this case we obtain

$$
\begin{aligned}
\lambda_{12} & =\frac{\left\langle\xi^{2}(\mathbf{h}) \xi\left[\mathbf{h}\left(\mathbf{I}-\mathbf{R}_{s}\right)\right]\right\rangle}{m^{3 / 2} p_{\mathbf{h}} \sqrt{p_{\mathbf{h}}\left(\mathbf{I}-\mathbf{R}_{s}\right)}} \\
& =\frac{\left\langle\sum_{1}^{m} \xi_{q}\left[\mathbf{h}\left(\mathbf{I}-\mathbf{R}_{q}\right)\right] \xi\left[\mathbf{h}\left(\mathbf{I}-\mathbf{R}_{s}\right)\right]\right\rangle}{m^{3 / 2} p_{\mathbf{h}} \sqrt{p_{\mathbf{h}}\left(\mathbf{I}-\mathbf{R}_{s}\right)}}=\frac{\sqrt{p_{\mathbf{h}}\left(\mathbf{I}-\mathbf{R}_{s}\right)}}{m^{1 / 2} p_{\mathbf{h}}}
\end{aligned}
$$

and

$$
\begin{aligned}
\left\langle E_{\mathbf{h}}\left(\mathbf{I}-\mathbf{R}_{s}\right) \mid E_{\mathbf{h}}\right\rangle & =\frac{1}{2 N^{1 / 2}} \frac{\sqrt{\left.p_{\mathbf{h}} \mathbf{I}-\mathbf{R}_{s}\right)}}{p_{\mathbf{h}}}\left(\left|E_{\mathbf{h}}\right|^{2}-1\right) \\
& \times \exp 2 \pi i \mathbf{h} T_{s} .
\end{aligned}
$$

Likewise equation (19) is modified to

$$
\begin{aligned}
P_{+}\left[E_{\mathbf{h}}\left(\mathbf{I}-\mathbf{R}_{s}\right)\right] & =\frac{1}{2}+\frac{1}{2} \tanh \frac{W}{2 N^{1 / 2}}\left|E_{\mathbf{h}}\left(\mathbf{I}-\mathbf{R}_{s}\right)\right|\left(\left|E_{\mathbf{h}}\right|^{2}-1\right) \\
& \times \exp 2 \pi i \mathbf{h} \mathbf{T}_{s}
\end{aligned}
$$

where $W$ is equal to $\left.\sqrt{p_{\mathbf{h}}(\mathbf{I}-} \overline{\mathbf{R}_{s}}\right) / p_{\mathbf{h}}$.

## Non-centrosymmetric crystal

As is well known, the characteristic function $C$ of the multivariate distribution $P\left(A_{1}, A_{2}, A_{3}, B_{1}, B_{2}, B_{3}\right)$ may be expanded in terms of cumulants:

$$
\begin{aligned}
C\left(u_{1}, u_{2}, u_{3}, v_{1}, v_{2}, v_{3}\right)= & \exp \left[t \sum_{2}^{\infty} r+s+\ldots+w \frac{\lambda_{r s \ldots w}^{\prime}}{r!s!\ldots w!}\right. \\
& \left.\times\left(\frac{i u_{1}}{t^{1 / 2}}\right)^{r}\left(\frac{i u_{2}}{t^{1 / 2}}\right)^{s} \cdots\left(\frac{i v_{3}}{t^{1 / 2}}\right)^{w}\right]
\end{aligned}
$$

where

$$
\lambda_{r s \ldots w}^{\prime}=\frac{K_{r s \ldots w}}{m^{(r+s+\ldots+w) / 2}}
$$

$K_{r s \ldots w}$ is a cumulant (with indices $r, s, \ldots, w$ ) of the distribution:

$$
\begin{aligned}
A_{1} & =\left|E_{\mathrm{h}}\right| \cos \varphi_{\mathrm{h}} \\
A_{2} & =\left|E_{\mathrm{k}}\right| \cos \varphi_{\mathrm{k}} \\
A_{3} & =\left|E_{\mathrm{h}-\mathrm{k}}\right| \cos \varphi_{\mathrm{h}-\mathrm{k}}, \ldots
\end{aligned}
$$

By taking the Fourier transform we can derive (retaining terms to order $1 / t^{1 / 2}$ ) the formula

$$
\begin{align*}
& P\left(A_{1}, A_{2}, A_{3}, B_{1}, B_{2}, B_{3}\right)=\frac{1}{(2 \pi)^{3}} \cdot \frac{1}{\sqrt{\lambda}} \\
& \quad \times \exp \left\{-\frac{1}{2}\left[\frac{A_{1}^{2}}{\lambda_{200000}^{\prime}}+\frac{A_{2}^{2}}{\lambda_{020000}^{\prime}}+\ldots+\frac{B_{2}^{2}}{\lambda_{000020}^{\prime}}\right.\right. \\
& \left.\left.\quad+\frac{B_{3}^{2}}{\lambda_{000002}^{\prime}}\right]\right\} \cdot\left\{1+\frac{1}{t^{1 / 2}}\left[\frac{\lambda_{111000}^{\prime} A_{1} A_{2} A_{3}}{\lambda_{200000}^{\prime} \cdot \lambda_{020000}^{\prime} \cdot \lambda_{002000}^{\prime}}\right.\right. \\
& \quad+\frac{\lambda_{001110}^{\prime} A_{3} B_{1} B_{2}}{\lambda_{002000}^{\prime} \cdot \lambda_{000200}^{\prime} \cdot \lambda_{000020}^{\prime}}+\frac{\lambda_{010101}^{\prime} A_{2} B_{1} B_{3}}{\lambda_{020000}^{\prime} \cdot \lambda_{000200}^{\prime} \cdot \lambda_{000002}^{\prime}} \\
& \left.\quad+\frac{\lambda_{100011}^{\prime} A_{1} B_{2} B_{3}}{\lambda_{200000}^{\prime} \cdot \lambda_{000020}^{\prime} \cdot \lambda_{000002}^{\prime}}+\ldots\right\}, \tag{20}
\end{align*}
$$

where

$$
\lambda=\lambda_{200000}^{\prime} \cdot \lambda_{020000}^{\prime} \cdots \lambda_{000002}^{\prime}
$$

and

$$
\lambda_{200000}^{\prime}=\frac{K_{200000}}{m}=\frac{\left\langle\psi^{2}(\mathbf{h})\right\rangle}{m p_{\mathbf{h}}}, \ldots
$$

It is easily shown that the distribution (20) coincides, in the case of general reflexions, with the known formula (Karle \& Hauptman, 1956),

$$
\begin{aligned}
& P\left(\left|E_{1}\right|,\left|E_{2}\right|,\left|E_{3}\right|, \varphi_{1}, \varphi_{2}, \varphi_{3}\right)=-\frac{1}{\pi^{3}}\left|E_{1}\right|\left|E_{2}\right|\left|E_{3}\right| \\
& \quad \times \exp \left(-\left|E_{1}\right|^{2}-\left|E_{2}\right|^{2}-\left|E_{3}\right|^{2}\right) \\
& \quad \times\left\{1+\frac{2}{V N}\left|E_{1}\right|\left|E_{2}\right|\left|E_{3}\right| \cos \left(\varphi_{1}-\varphi_{2}-\varphi_{3}\right)\right\} .
\end{aligned}
$$

The conditional mean values

$$
\left\langle A_{2} A_{3}-B_{2} B_{3}\right\rangle=\langle | E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\left|\cos \left(\varphi_{\mathbf{k}}+\varphi_{\mathbf{h}-\mathbf{k}}\right)\right\rangle,
$$

and

$$
\left\langle A_{2} B_{3}+A_{3} B_{2}\right\rangle=\langle | E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\left|\sin \left(\varphi_{\mathbf{k}}+\varphi_{\mathbf{h}-\mathbf{k}}\right)\right\rangle,
$$

when $A_{1}$ and $B_{1}$ are fixed and the fact that

$$
K_{111000}=m_{111000}=\frac{\langle\psi(\mathbf{h}) \psi(\mathbf{k}) \psi(\mathbf{h}-\mathbf{k})\rangle}{\sqrt{p_{\mathbf{h}} p_{\mathbf{k}} p_{\mathbf{h}-\mathbf{k}}}}, \text { etc. }
$$

gives the result

$$
\begin{align*}
& \langle | E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\left|\cos \left(\varphi_{\mathbf{k}}+\varphi_{\mathbf{h} \mathbf{- k}}\right)\right\rangle \\
& \quad=\frac{1}{t^{1 / 2}}\left\{\frac{\lambda_{11000}^{\prime}-\lambda_{100011}^{\prime}}{\lambda_{20000}^{\prime}} A_{1}\right\}=\frac{1}{N^{1 / 2}} \frac{\sqrt{p_{\mathbf{h}}}}{\left\langle\psi^{2}(\mathbf{h})\right\rangle} \\
& \quad \times\left\{\frac{\langle\psi(\mathbf{h})[\psi(\mathbf{k}) \psi(\mathbf{h}-\mathbf{k})-\eta(\mathbf{k}) \eta(\mathbf{h}-\mathbf{k})]\rangle}{\sqrt{p_{\mathbf{k}} p_{\mathbf{h}-\mathbf{k}}}}\right\} \\
& \quad \times\left|E_{\mathbf{h}}\right| \cos \varphi_{\mathbf{h}} . \tag{21}
\end{align*}
$$

In the same way we find

$$
\begin{align*}
& \langle | E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\left|\sin \left(\varphi_{\mathbf{k}}+\varphi_{\mathbf{h}-\mathbf{k}}\right)\right\rangle \\
& \quad=\frac{1}{t^{1 / 2}}\left\{\frac{\lambda_{001110}^{\prime}+\lambda_{010101}^{\prime}}{\lambda_{000200}^{\prime}} B_{1}\right\}=\frac{1}{N^{1 / 2}} \frac{\sqrt{p_{\mathbf{h}}}}{\left\langle\eta^{2}(\mathbf{h})\right\rangle} \\
& \quad \times\left\{\frac{\langle\eta(\mathbf{h})[\eta(\mathbf{k}) \psi(\mathbf{h}-\mathbf{k})+\psi(\mathbf{k}) \eta(\mathbf{h}-\mathbf{k})]\rangle}{\sqrt{p_{\mathbf{k}} p_{\mathbf{h}-\mathbf{k}}}}\right\} \\
& \quad \times\left|E_{\mathbf{h}}\right| \sin \varphi_{\mathbf{h}} . \tag{22}
\end{align*}
$$

From the distribution (20) we derive the conditional mean values $\left\langle A_{\mathbf{h}}\right\rangle$ and $\left\langle B_{\mathbf{h}}\right\rangle$ when $A_{\mathbf{k}}, A_{\mathbf{h}-\mathbf{k}}, B_{\mathbf{k}}, B_{\mathbf{h}-\mathbf{k}}$ are fixed. After some calculations

$$
\begin{aligned}
& \langle | E_{\mathbf{h}}\left|\cos \varphi_{\mathbf{h}}\right\rangle=\frac{1}{t^{1 / 2}}\left\{\lambda_{11000}^{\prime} \frac{A_{2} A_{3}}{\lambda_{020000}^{\prime} \cdot \lambda_{002000}^{\prime}}\right. \\
& \quad+\frac{\lambda_{100011}^{\prime}}{\left.\lambda_{000020}^{\prime} \cdot \lambda_{000002}^{\prime}-B_{2} B_{3}\right\}} \\
& \quad=\frac{\sqrt{p_{\mathbf{k}} p_{\mathbf{h}-\mathbf{k}}}}{\sqrt{p_{\mathbf{h}}}} \frac{m}{N^{1 / 2}}\left|E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\right|\{\langle\psi(\mathbf{h}) \psi(\mathbf{k}) \psi(\mathbf{h}-\mathbf{k})\rangle \\
& \quad \times \cos \varphi_{\mathbf{k}} \cos \varphi_{\mathbf{h}-\mathbf{k}} \\
& \left.\quad+\frac{\langle\psi(\mathbf{h}) \eta(\mathbf{k}) \eta(\mathbf{h}-\mathbf{k})\rangle}{\left\langle\eta^{2}(\mathbf{k})\right\rangle\left\langle\eta^{2}(\mathbf{h}-\mathbf{k})\right\rangle} \sin \varphi_{\mathbf{k}} \sin \varphi_{\mathbf{h}-\mathbf{k}}\right\} ; \\
& \langle | E_{\mathbf{h}}\left|\sin \varphi_{\mathbf{h}}\right\rangle=\frac{\sqrt{\left.\left.p_{\mathbf{k}} p_{\mathbf{h}}-\mathbf{k}\right)\right\rangle}}{\sqrt{p_{\mathbf{h}}}} \frac{m}{N^{1 / 2}}\left|E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\right|
\end{aligned}
$$

$$
\begin{align*}
& \times\left\{\frac{\langle\eta(\mathbf{h}) \eta(\mathbf{k}) \psi(\mathbf{h}-\mathbf{k})\rangle}{\left\langle\eta^{2}(\mathbf{k})\right\rangle\left\langle\psi^{2}(\mathbf{h}-\mathbf{k})\right\rangle} \sin \varphi_{\mathbf{k}} \cos \varphi_{\mathbf{h}-\mathbf{k}}\right. \\
& \left.+\frac{\langle\eta(\mathbf{h}) \psi(\mathbf{k}) \eta(\mathbf{h}-\mathbf{k})\rangle}{\left\langle\psi^{2}(\mathbf{k})\right\rangle\left\langle\eta^{2}(\mathbf{h}-\mathbf{k})\right\rangle} \cos \varphi_{\mathbf{k}} \sin \varphi_{\mathbf{h}-\mathbf{k}}\right\} . \tag{24}
\end{align*}
$$

If $E_{\mathbf{h}}, E_{\mathbf{k}}, E_{\mathbf{h}-\mathbf{k}}$ are general reflexions we obtain

$$
\begin{aligned}
& \langle | E_{\mathbf{h}}\left|\cos \varphi_{\mathbf{h}}\right\rangle=\frac{1}{V N}\left|E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\right| \cos \left(\varphi_{\mathbf{k}}+\varphi_{\mathbf{h}-\mathbf{k}}\right), \\
& \langle | E_{\mathbf{h}}\left|\sin \varphi_{\mathbf{h}}\right\rangle=\frac{1}{V N}\left|E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\right| \sin \left(\varphi_{\mathbf{k}}+\varphi_{\mathbf{h}-\mathbf{k}}\right)
\end{aligned}
$$

so that in all space groups the relation (8) is justified. If $E_{\mathbf{k}}$ is a centrosymmetric reflexion [ $\eta(\mathbf{k})=0$ ], we find

$$
\begin{align*}
& \langle | E_{\mathbf{h}}\left|\cos \varphi_{\mathbf{h}}\right\rangle=\frac{\sqrt{p_{\mathbf{k}} p_{\mathbf{h}-\mathbf{k}}}}{\sqrt{p_{\mathbf{h}}}} \frac{m}{N^{1 / 2}}\left|E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\right| \\
& \quad \times\left\{\frac{\langle\psi(\mathbf{h}) \psi(\mathbf{k}) \psi(\mathbf{h}-\mathbf{k})\rangle}{\left\langle\psi^{2}(\mathbf{k})\right\rangle\left\langle\psi^{2}(\mathbf{h}-\mathbf{k})\right\rangle} \cos \left(\varphi_{\mathbf{k}}+\varphi_{\mathbf{h}-\mathbf{k}}\right)\right\},  \tag{25}\\
& \langle | E_{\mathbf{h}}\left|\sin \varphi_{\mathbf{h}}\right\rangle=\frac{\sqrt{p_{\mathbf{k}} p_{\mathbf{h}-\mathbf{k}}}}{\sqrt{p_{\mathbf{h}}}} \frac{m}{N^{1 / 2}}\left|E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\right| \\
& \quad \times\left\{\frac{\langle\eta(\mathbf{h}) \psi(\mathbf{k}) \eta(\mathbf{h}-\mathbf{k})\rangle}{\left\langle\psi^{2}(\mathbf{k})\right\rangle\left\langle\psi^{2}(\mathbf{h}-\mathbf{k})\right\rangle} \sin \left(\varphi_{\mathbf{k}}+\varphi_{\mathbf{h}-\mathbf{k}}\right)\right\} . \tag{26}
\end{align*}
$$

As in this case

$$
\begin{equation*}
\frac{\langle\psi(\mathbf{h}) \psi(\mathbf{k}) \psi(\mathbf{h}-\mathbf{k})\rangle}{\left\langle\psi^{2}(\mathbf{k})\right\rangle\left\langle\psi^{2}(\mathbf{h}-\mathbf{k})\right\rangle}=\frac{\langle\eta(\mathbf{h}) \psi(\mathbf{k}) \eta(\mathbf{h}-\mathbf{k})\rangle}{\left\langle\psi^{2}(\mathbf{k})\right\rangle\left\langle\psi^{2}(\mathbf{h}-\mathbf{k})\right\rangle} \tag{27}
\end{equation*}
$$

the relation (8) is still valid.
Analogously, if one $E_{\mathbf{h}-\mathbf{k}}$ reflexion is centrosymmetrical, equation (25) still holds, and relations similar to (26) and (27) can be worked out. Following Cochran (1955), one can easily deduce that in the distribution (9) a suitable weight must be applied: in the example of Table 1,

$$
\begin{align*}
G_{\mathbf{h}, \mathbf{k}} & =m \frac{\sqrt{p_{\mathbf{k}} p_{\mathbf{h}}-\mathbf{k}}}{\sqrt{p_{\mathbf{h}}}}\left\langle\frac{\langle(\mathbf{h}) \psi(\mathbf{k}) \psi(\mathbf{h}-\mathbf{k})\rangle}{\left\langle\psi^{2}(\mathbf{k})\right\rangle\left\langle\psi^{2}(\mathbf{h}-\mathbf{k})\right\rangle}\right. \\
& \times 2 \frac{\left|E_{\mathbf{h}} E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{- k}}\right|}{V N}=W_{\mathbf{h}, \mathbf{k}} 2 \frac{\left|E_{\mathbf{h}} E_{\mathbf{k}} E_{\mathbf{h}-\mathbf{k}}\right|}{V N} . \tag{28}
\end{align*}
$$

A more general expression for these weights, valid in all space groups will be identified in the following paper (Giacovazzo, 1974).

I wish to thank Dr J. Karle for critical reading of the manuscript.

## APPENDIX

From the theory of linearization (Bertaut, 1959a,b) we obtain for a general space group of order $m$,

|  | $h_{1} k_{1} l_{1}$ | $h_{1} k_{1} l_{1}$ | $h_{1} k_{1} l_{1}$ | $h_{1} k_{1} l_{1}$ | $h_{1} k_{1} l_{1}$ | $h_{1} k_{1} 0$ | $h_{1} k_{1} 0$ | $h_{1} k_{1} 0$ | $h_{1} k_{1} 0$ | $h_{1} 00$ | $h_{1} 00$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $h_{2} k_{2} l_{2}$ | $h_{2} k_{2} 0$ | $h_{2} 00$ | $h_{2} k_{2} 0$ | $h_{2} 00$ | $h_{2} k_{2} 0$ | $h_{2} 0 l_{2}$ | $h_{2} k_{2} 0$ | $h_{2} 00$ | $h_{2} 00$ | $h_{2} k_{2} l_{2}$ |
|  | $h_{3} k_{3} l_{3}$ | $h_{3} k_{3} l_{3}$ | $h_{3} k_{3} l_{3}$ | $h_{3} 0 l_{3}$ | $0 k_{3} l_{3}$ | $h_{3} k_{3} 0$ | $0 k_{3} l_{3}$ | $h_{3} 00$ | $0 k_{3} 0$ | $h_{3} 00$ | $h_{3} k_{3} l_{3}$ |
| $\langle\psi(\mathbf{h}) \psi(\mathbf{k}) \psi(\mathbf{h}-\mathbf{k})\rangle$ | 1 | 2 | 4 | 4 | 8 | 4 | 8 | 8 | 16 | 16 | 4 |
| $\langle\eta(\mathbf{h}) \eta(\mathbf{k}) \psi(\mathbf{h}-\mathbf{k})\rangle$ | 1 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| $\langle\eta(\mathbf{h}) \psi(\mathbf{k}) \eta(\mathbf{h}-\mathbf{k})\rangle$ | 1 | 2 | 4 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| $\langle\psi(\mathbf{h}) \eta(\mathbf{k}) \eta(\mathbf{h}-\mathbf{k})\rangle$ | -1 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | -4 |
| $W_{\mathbf{h}, \mathbf{k}}$ | 1 | 1 | $\checkmark 2$ | 1 | $\checkmark 2$ | 1 | 2 | $\checkmark 2$ | 2 | $\checkmark 2$ | $2 / 2$ |

Table 2. $\mathbf{h}_{\boldsymbol{i}}+\mathbf{h}_{j}+\mathbf{h}_{k}=\mathbf{0}$

| Parity | $h_{1} k_{1} l_{1}$ | $h_{1} k_{1} l_{1}$ | $h_{1} k_{1} l_{1}$ | $h_{1} k_{1} l_{1}$ | $h_{1} k_{1} l_{1}$ | $h_{1} k_{1} 0$ | $h_{1} k_{1} 0$ | $h_{1} k_{1} 0$ | $h_{1} k_{1} 0$ | $h_{1} 00$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| classes | $h_{2} k_{2} l_{2}$ | $h_{2} k_{2} 0$ | $h_{2} 00$ | $h_{2} k_{2} 0$ | $h_{2} 00$ | $h_{2} k_{2} 0$ | $h_{2} 0 l_{2}$ | $h_{2} k_{2} 0$ | $h_{2} 00$ | $h_{2} 00$ |
|  | $h_{3} k_{3} l_{3}$ | $h_{3} k_{3} l_{3}$ | $h_{3} k_{3} l_{3}$ | $h_{3} 0 l_{3}$ | $0 k_{3} l_{3}$ | $h_{3} k_{3} 0$ | $0 k_{3} l_{3}$ | $h_{3} 00$ | $0 k_{3} 0$ | $h_{3} 00$ |
| $\langle\xi(\mathbf{h}) \xi(\mathbf{k}) \xi(\mathbf{h}+\mathbf{k})\rangle$ | 8 | 16 | 32 | 32 | 64 | 32 | 64 | 64 | 128 | 128 |
| $\underline{\langle\xi(\mathbf{h}) \xi(\mathbf{k}) \xi(\mathbf{h}+\mathbf{k})\rangle}$ |  |  |  |  |  |  |  |  |  |  |
| $\overline{\sqrt{p_{\mathbf{h}}}} / \overline{\overline{p_{\mathbf{k}}}} / \sqrt{\overline{p_{\mathbf{h}+\mathbf{k}}}}$ | 8 | 8/2 | 16 | 16 | 16/2 | 8/2 | 16/2 | 16 | 16/2 | 16 |
| $W_{\mathrm{h}, \mathbf{k}}$ | 1 | $\sqrt{ } 2$ | 2 | 2 | $2 \sqrt{2}$ | $\checkmark 2$ | $2 \gamma / 2$ | 2 | $2 / 2$ | 2 |

$$
\begin{align*}
\xi\left(\mathbf{H}_{3}\right) \xi\left(\mathbf{H}_{1}\right)=\sum_{1}^{m} \xi\left(\mathbf{H}_{3}\right. & \left.+\mathbf{H}_{1} \mathbf{C}_{s}\right) \\
& =\sum_{1}^{m} a_{s}\left(\mathbf{H}_{1}\right) \xi\left(\mathbf{H}_{3}+\mathbf{H}_{1} \mathbf{R}_{s}\right) \tag{Al}
\end{align*}
$$

where $a_{s}(\mathbf{H})=\exp 2 \pi i \mathbf{H T}$.
If we multiply equation (A1) for $\xi\left(\mathbf{H}_{2}\right)$, by setting $\mathbf{H}_{3}=\overline{\mathbf{H}}_{1}+\overline{\mathbf{H}}_{2}$, we find

$$
\begin{align*}
& \xi\left(\mathbf{H}_{1}\right) \xi\left(\mathbf{H}_{2}\right) \xi\left(\mathbf{H}_{3}\right)=\sum_{1}^{m} \sum_{1}^{m} \xi\left[\mathbf{H}_{\mathbf{1}}\left(\mathbf{C}_{s}-\mathbf{I}\right)+\mathbf{H}_{2}\left(\mathbf{C}_{r}-\mathbf{I}\right)\right] \\
&=\sum_{1}^{m} s, r \\
& \times \xi\left[\mathbf{H}_{\mathbf{1}}\left(\mathbf{H}_{1}\right) a_{r}\left(\mathbf{H}_{2}\right)\right.  \tag{A2}\\
&\left.\mathbf{I})+\mathbf{H}_{2}\left(\mathbf{R}_{r}-\mathbf{I}\right)\right]
\end{align*}
$$

The mean value $\left\langle\xi\left(\mathbf{H}_{1}\right) \xi\left(\mathbf{H}_{2}\right) \xi\left(\mathbf{H}_{3}\right)\right\rangle$ is different from the zero for all $\mathbf{C}_{r}, \mathbf{C}_{s}$ operations for which

$$
\begin{equation*}
\mathbf{H}_{1}\left(\mathbf{R}_{s}-\mathbf{I}\right)+\mathbf{H}_{2}\left(\mathbf{R}_{r}-\mathbf{I}\right)=0 \tag{A3}
\end{equation*}
$$

For example, if $\mathbf{C}_{s}$ is a operation for which $\mathbf{H}_{1}\left(\mathbf{R}_{s}-\mathbf{I}\right)=0$, the condition (A3) is verified for all $r$ operations $\mathbf{C}_{r}$ such that

$$
\mathbf{H}_{2}\left(\mathbf{R}_{r}-\mathbf{I}\right)=0
$$

Therefore, if $E_{\mathbf{H}_{3}}$ is non-special reflexion, we obtain

$$
\left\langle\xi\left(\mathbf{H}_{1}\right) \xi\left(\mathbf{H}_{2}\right) \xi\left(\mathbf{H}_{3}\right)\right\rangle=p_{\mathbf{H}_{1}} p_{\mathbf{H}_{2}} \xi(0)=p_{\mathbf{H}_{1}} p_{\mathbf{H}_{2}} m .
$$

Numerical values for different parity classes are shown in Table 2 for the space group Pmmm.
In a similar way it results

$$
\xi^{2}(\mathbf{h}) \xi(2 \mathbf{h})=\sum_{1}^{m} s, r\left[\mathbf{h}\left(\mathbf{I}-\mathbf{C}_{s}+2 \mathbf{C}_{r}\right)\right]
$$

If the $\mathbf{h}$ reflexion is general, $\left\langle\xi^{2}(\mathbf{h}) \xi(2 \mathbf{h})\right\rangle$ is different from zero for $\mathbf{C}_{s}=-\mathbf{I}$ and $\mathbf{C}_{r}=\mathbf{I}$ : then

$$
\left\langle\xi^{2}(\mathbf{h}) \xi(2 \mathbf{h})\right\rangle=m .
$$

If $E_{\mathbf{h}}$ has statistical weight $p_{\mathrm{h}}$,

$$
\left\langle\xi^{2}(\mathbf{h}) \xi(2 \mathbf{h})\right\rangle=p_{\mathbf{h}} \sum_{1}^{m / p \mathbf{h}} s, r\left[\mathbf{h}\left(\mathbf{I}-\mathbf{C}_{s}+2 \mathbf{C}_{r}\right)\right]=m p_{\mathbf{h}}
$$

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